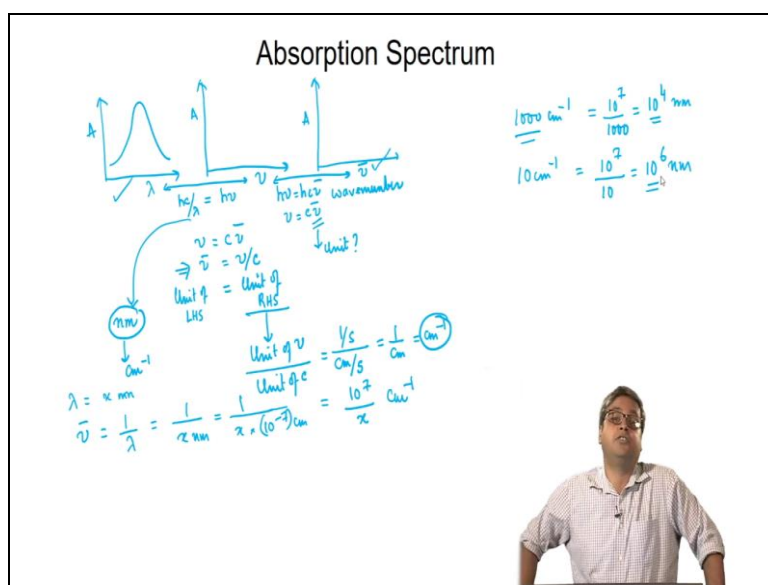


Fundamentals of Spectroscopy
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Lecture – 9
Lineshape Analysis

Hello everyone welcome to this lecture in the last lecture we have discussed how we can obtain an absorption spectrum.

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We had mentioned that if you plot absorbance obtained from a spectroscopy experiment against wavelength lambda we can obtain an absorption spectrum. Although conventionally absorbance is plotted against wavelength we can also plot absorbance against frequency as these are interchangeable. So, we can write $h\nu$ equals hc/λ . Moreover we can also plot absorbance against wave number.

Wave number is denoted by $\bar{\nu}$ a new bar is related to frequency because $h\nu$ equals $hc\bar{\nu}$ or we can write ν equals $c\bar{\nu}$. So, the next question is what is the unit of $\bar{\nu}$ because we can write ν equals $c\bar{\nu}$ or rearranging this equation we can write $\bar{\nu}$ equals ν/c . So, we know the unit of the left hand side is equal to the unit of the right hand side. So, in order to find the unit of $\bar{\nu}$ we have to find the unit of the right hand side and the unit of the right hand side is unit of ν by unit of c that is $1/\text{second}$ by $\text{centimeter per second}$ this gives $1/\text{centimeter}$ or $\text{centimeter inverse}$.

So, centimeter inverse is the unit of nu bar and also lambda is normally represented in the unit of nanometers. So, let us look into the relation between nanometer and centimeter inverse. So, let us say the wavelength lambda equals x nanometer, so in centimeter inverse the wave number will be 1 by lambda that is 1 by x nanometer that is 1 by x times 10 to the power - 7 centimeter equals 10 to the power 7 by x centimeter inverse.

Later we will discuss about different forms of spectroscopy like UV visible spectroscopy infrared spectroscopy microwave spectroscopy and we will see that we get different spectroscopy depending on the wavelength of the light or the frequency of the light interacting with the matter. Conventionally when we plot our UV visible spectrum we use wavelength in the x axis however when we plot an infrared spectrum conventionally we plot wave numbers on the x axis this is just a convention used for convenience such that the values plotted on the x axis are not very large numbers.

So, let us see what do I mean by that for example 1000 wave numbers is a typical higher frequency so if we convert this number into nanometer it will be 10 to the power 7 by 1000 equals 10 to the power 4 nanometer, so 10 to the power 4 is a larger number than 1000. Similarly 10 wave numbers is a typical energy gap involved in a microwave transition, so this amounts to 10 to the power 7 divided by 10 that is 10 to the power 6 nanometer which is a very, very large number.

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Spectral Width

$E = h\nu_{12}$
 $\Delta E = E_2 - E_1 = h\nu_{12}$
 $A = \log \frac{I_0}{I_f} \quad I_f = I_0$
 $= \log \frac{I_0}{I_1} = \log 1 = 0$

Lorentzian function
 $f(\nu - \nu_0)$

Transmission
Wavenumber / cm^{-1}
 CO_2 Lorentzian
2368.3 2369.1 Wavenumber / cm^{-1}

So, far we have explained the absorption process by saying we have a molecule in a stationary state of energy E_1 and when light of energy equals $h\nu_{12}$ interacts with the molecule then the molecule goes from the stationary state with energy E_1 to another stationary state of energy E_2 such that Δe or $E_2 - E_1$ equals $h\nu_{12}$ apparently only one frequency is involved that is ν_{12} in the transition process or in other words in terms of experiments if we use a polychromatic light source that is light having many, many frequencies we can expect that only one frequency that is ν_{12} should be absorbed by the matter.

So, in all other frequencies where light is not absorbed absorbance is given by A equals $\log I_0$ by I_T and because in the other frequencies light is not absorbed by the matter so I_T for those frequencies will be equal to I_0 , so absorbance we can write $\log I_0$ divided by I_0 equals $\log 1$ that is 0. The absorbance should only be non0 at the frequency ν_{12} so this can be represented in a spectrum so again we are plotting absorbance versus frequency and let us say this is my frequency ν_{12} .

So for all other frequencies the absorbance will be 0 and only at ν_{12} the absorbance will be non0 however it has been seen from experiments that the spectroscopic lines are not as infinitely sharp as shown in the plot. But they are more or less broad additionally we should realize that as the light sources polychromatic that is the light consists of many frequencies thus multiple transitions can be obtained from the same sample.

So, this is a spectrum where the lines are associated with the anti symmetric stretching mode of carbon dioxide. At high resolution the spectrum seems to consist of multiple lines from multiple transitions so if we expand the scale or if we zoom into one particular line the line with a per N narrow feature is observed to have a definite width and a characteristic shape. The shape of the line spectrum is known as line shape.

As this shape is a function of frequency the line shape can be represented using a line shape function which is given by $f(\nu - \nu_0)$ where ν_0 is the frequency corresponding to the maximum absorbance. So, ν_0 is also known as the peak position for a single peak corresponding to a transition the shape of this spectrum is symmetric with respect to ν_0 . So, the observation of the width and line shape from the carbon-dioxide spectrum immediately raises two questions.

One what are the possible functional forms of these shapes and number 2 what physical processes are responsible for these shapes.

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Line Broadening - Classification

✓ (1) homogeneous
(2) inhomogeneous

$[M] = 0.1 \text{ M}$

1 ml

For 1000 ml of sample, concentration = 1M
 $N_{Av} \approx 10^{23}$

$$\frac{10^{23}}{10 \times 1000} = 10^{19} \text{ molecules}$$

Line shape functions fall into 1 of the 2 general categories one is homogeneous and the second is in homogeneous. A homogeneous line shape occurs when all the molecules in the system have identical line shape functions. Here we should understand that we have an ensemble of molecules in the sample solution on which we shine light during the spectroscopy experiment. So, let us say we have made a sample solution of concentration equal to 0.1 molar and let us say the volume of the container which is placed in the direction of light so that the light is passed to the container and goes through the detector.

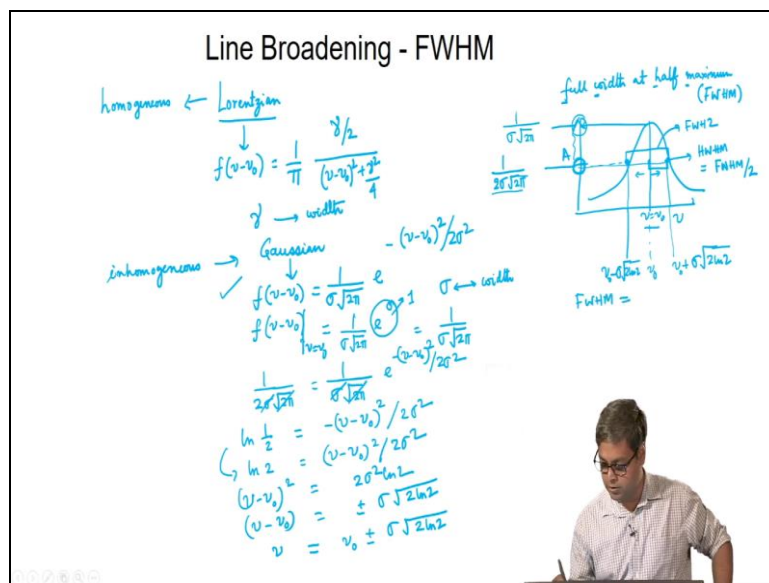
So, the volume of this container is 1 milliliter so we know that for 1000 milliliter of sample with concentration equal to 1 molar we have a Avogadro number of molecules which is approximately equal to 10 to the power 23 molecules. But now in our solution we have a volume of 1 milliliter and a concentration of 0.1 molar so the number of molecules we have during the experiment is 10 to the power 23 divided by 10 times 1000 equals 10 to the power 19 molecules.

Thus if all the molecules are affected the same way all our homogeneously affected during the light matter interaction they will have the same line shapes and we will have homogeneous broadening or homogeneous line shape function. For example if the absorbing species in the gas phase is subjected to high pressure then all the molecules are found to have an identical

pressure broadened line shape for a particular transition pressure broadening of a transition is thus say to be a homogeneous broadening.

In contrast if the sample is dissolved in a liquid solvent then this order inherent in the structure of the liquid provides different solvent environments for the solute. As each solute experiences a slightly different environment they have a slightly different spectrum. The observed absorption spectrum from the experiment is made up of all the different spectra for the different molecular environments and it is said to be inhomogeneously broadened.

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In general the homogeneously broadened spectrum can be represented by a Lorentzian. So, the Lorentzian line shape function represents homogeneously broadened spectrum. The Lorentzian line shape function is given by $f(\nu - \nu_0) = \frac{1}{\pi} \frac{\gamma}{(\nu - \nu_0)^2 + \frac{\gamma^2}{4}}$ so here γ is the parameter specifying the width of the spectrum. The inhomogeneously broadened spectrum is given by a Gaussian line shape function.

And the Gaussian line shape function can be written as $f(\nu - \nu_0) = \frac{1}{\sigma\sqrt{2\pi}} e^{-\frac{(\nu - \nu_0)^2}{2\sigma^2}}$, here σ is related to the width of the spectrum. So, we can quantify width of the observed spectrum in terms of full width at half maximum. This is also written as FWHM from the initials of full width at half maximum. So, let us try to understand how to obtain full width half maximum or FWHM from a spectrum.

So let us say we have an absorption spectrum so we are plotting absorbance versus frequency. So, first thing we have to find out what is the absorbance at $\nu = \nu_0$ because at ν_0 the absorbance is maximum. Once we find the maximum value of absorbance we have to find what is half of that maximum value and once we find that half of the maximum value of absorbance we have to find the values of the frequency that corresponds to half of the values of the absorbance.

So because we are considering a spectrum which is symmetric about ν_0 , so we will have two such frequencies one on the left and one on the right of your 0. So, now the frequency difference between these two points is the full width half maximum. Now instead of taking this frequency difference if we take the frequency difference between ν_0 and one of these frequencies which has half the maximum absorbance then that is given by half width at half maximum and because the spectrum is symmetric half width at half maximum equals full width at half maximum by 2.

So, if the functional form of the line shape is known we can find the functional form of the full width at half maximum. So, let us look into the Gaussian functional form. So, we have to evaluate this function at $\nu = \nu_0$ so at $\nu = \nu_0$ $f(\nu - \nu_0)$ becomes $1 / (\sigma \sqrt{2\pi}) e^{-0}$ because e^0 is 1 so this is $1 / (\sigma \sqrt{2\pi})$. So, now that means the maximum absorbance is $1 / (\sigma \sqrt{2\pi})$.

So, the absorbance at half the maximum absorbance value is $1 / (2\sigma \sqrt{2\pi})$. So, now we have to find the frequency at which the absorbance is $1 / (2\sigma \sqrt{2\pi})$, so we can write $1 / (2\sigma \sqrt{2\pi}) = 1 / (\sigma \sqrt{2\pi}) e^{-\frac{(\nu - \nu_0)^2}{2\sigma^2}}$. So, we will cancel out this so we have or we can write $\ln \frac{1}{2} = -\frac{(\nu - \nu_0)^2}{2\sigma^2}$.

So, we can also write from this equation $\ln 2 = \frac{(\nu - \nu_0)^2}{2\sigma^2}$ or $(\nu - \nu_0)^2 = 2\sigma^2 \ln 2$ or $\nu - \nu_0 = \pm \sigma \sqrt{2 \ln 2}$. So, now we needed to find ν , so $\nu = \nu_0 \pm \sigma \sqrt{2 \ln 2}$. So what does it mean so I have ν_0 so this frequency is $\nu_0 + \sigma \sqrt{2 \ln 2}$ and this is $\nu_0 - \sigma \sqrt{2 \ln 2}$ so the full width half maximum equals to $2\sigma \sqrt{2 \ln 2}$.

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Line Broadening – Mechanisms Doppler broadening

$$\Delta E \Delta t \approx \hbar$$

$$\Delta E \approx \frac{\hbar}{\Delta t} = \frac{h}{2\pi \Delta t}$$

$$\Delta E = \frac{h}{2\pi \tau}$$

$$\hbar \Delta \nu = \frac{h}{2\pi \tau}$$

$$\Delta \nu = \frac{1}{2\pi \tau}$$

$$\Delta \nu \propto \frac{1}{\tau}$$

$$A \rightarrow s^{-1}$$

$$A = \frac{1}{\tau}$$

$\nu = \nu_0 \left(1 + \frac{v}{c}\right)$
 $\Delta \nu = \frac{2\nu_0}{c} \sqrt{\frac{2 \ln 2 kT}{m}}$
 1 GHz
 $\text{H}_2 \sim 30 \text{ GHz}$

So, now let us look into the physical processes that lead to broadening of the spectrum one of them is due to quantum mechanical effects this arises from the Heisenberg's uncertainty principle between energy and time which is given by $\Delta E \Delta T$ is of the order of \hbar cross, so ΔE is the uncertainty in energy and ΔT is the uncertainty in time. The expression tells us that if the system survives in a quantum state for a finite time the energy of the state in principle cannot be known with accuracy better than ΔE which is given by \hbar cross by ΔT or \hbar by $2\pi \Delta T$.

Now let us think about spontaneous emission when we excite a molecule to a higher energy state Einstein hypothesized that the molecule will spontaneously decay from the higher energy state. If the molecule on average say decays after time τ so τ is the time the molecule survives in the excited state then τ is known as a lifetime and we can write ΔE equals \hbar by $2\pi \tau$ and for ΔE we can write $\hbar \Delta \nu$ equals \hbar by $2\pi \tau$ can cancel out \hbar so $\Delta \nu$ equals 1 by $2\pi \tau$.

So here $\Delta \nu$ is the uncertainty in frequency which gives rise to a broadening of the spectrum along the frequency axis. Thus the spectral broadening or $\Delta \nu$ is inversely proportional to the lifetime if the life time is short the spectrum will be broad and vice versa we have discussed that A is Einsteins coefficient for spontaneous emission. The unit of A is second inverse that is inverse of the unit of time.

It can be shown that E is equal to 1 by lifetime or 1 by τ . Moreover lifetime can be affected by interactions between the quantum state and the surrounding particles that is collisions. In the

condensed phase the occurrence of collisions are greater than that in the gas phase. As these collisions are inelastic in nature they reduce the lifetime of the excited state. A similar case happens when pressure is increased increase in pressure increases the number of collisions and thus reduces the lifetime.

In all these cases the spectral line shape is Lorentian and thus these factors lead to homogeneous broadening. One of the most common examples of gas phase inhomogeneous broadening occurs due to Maxwell Boltzmann distribution of molecular velocities and is called Doppler broadening. When viewed from the frame of the atom the Doppler effect results in light energy shift or shift in the frequency of light when the source is moving either toward or away from the observer.

When a source emitting radiation with frequency ν_0 moves with a velocity V and the observer is placed in the Z direction that is the detector is in the Z direction then the observer detects the radiation with a frequency not equal to ν_0 but ν which is equal to $\nu_0 (1 \pm \frac{V_z}{c})$ but this plus and minus sign is related to an approaching or receding source. It can be shown that the full width half maximum $\Delta \nu$ equals $2 \nu_0 \frac{V_z}{c} \sqrt{2 \ln 2} \sqrt{\frac{k T}{M}}$ where T is the temperature divided by M .

So we see that $\Delta \nu$ or full width half maximum of the Doppler broadening has an expression similar to what we had obtained for a Gaussian line shape. Thus Doppler broadening is inhomogeneous broadening and gives rise to a Gaussian line shape. We can see that the Doppler width is proportional to the square root of temperature and inversely proportional to the square root of mass. For example the Doppler width is 1 gigahertz for a UV visible transition at room temperature.

However for hydrogen the mass is small and the Doppler width is approximately much bigger that is approximately 30 gigahertz. So, finally I would like to end by mentioning that all these processes the Doppler effect the collisions the quantum mechanical effect that is lifetime happens simultaneously and thus the spectral line shape is never perfectly Gaussian or Lorentzian but is a convolution of Gaussian and Lorentzian line shape function and is given by the Voigt function or the profile of the spectrum is known as Voigt profile.