

Introduction to interfacial waves
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Lecture - 61
Derivation of the Stokes travelling wave (contd..)

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$$O(\epsilon^3): \quad \nabla^2 \phi_3 = 0 \quad \downarrow$$

$$\left(\frac{\partial \phi_3}{\partial z} \right)_0 + \eta_3 = - \left[\eta^2 \left(\frac{\partial^2 \phi_1}{\partial z^2 \partial \tau} \right)_0 + \frac{\eta^2}{2} \left(\frac{\partial^3 \phi_1}{\partial z^2 \partial \tau} \right)_0 + \eta_1 \left(\frac{\partial^2 \phi_2}{\partial z \partial \tau} \right)_0 \right.$$

$$+ \eta_1 \left(\frac{\partial \phi_1}{\partial x} \right)_0 \left(\frac{\partial^2 \phi_1}{\partial x \partial z} \right)_0 + \eta_1 \left(\frac{\partial \phi_1}{\partial y} \right)_0 \left(\frac{\partial^2 \phi_1}{\partial y \partial z} \right)_0$$

$$+ \left(\frac{\partial \phi_1}{\partial x} \right)_0 \left(\frac{\partial \phi_2}{\partial x} \right)_0 + \left(\frac{\partial \phi_1}{\partial y} \right)_0 \left(\frac{\partial \phi_2}{\partial y} \right)_0 + \left. \omega_2 \left(\frac{\partial \phi_1}{\partial z} \right)_0 \right]$$

→ Taylor series of the

$$\left(\frac{\partial \phi}{\partial t} \right)_{z=\eta} = (1 + \epsilon^2 \omega_2 + \dots) \frac{\partial}{\partial \tau} \left[\epsilon \phi_1 + \epsilon^2 \phi_2 \right]_{z=\eta}$$

$$= (1 + \epsilon^2 \omega_2) \left[\epsilon \left(\frac{\partial \phi_1}{\partial \tau} \right)_0 + \epsilon \left(\frac{\partial^2 \phi_1}{\partial z \partial \tau} \right)_0 (\epsilon \eta_1 + \epsilon^2 \eta_2) + \frac{\epsilon}{2} \left(\frac{\partial^3 \phi_1}{\partial z^2 \partial \tau} \right)_0 (\epsilon \eta_1)^2 \right]$$

In order to solve our Stokes wave problem up to order epsilon square we had to actually go to order epsilon cube this was necessary because we wanted to determine omega 2 the correction to the dispersion relation and this omega 2 did not appear in the equations up to order epsilon square. So, now we have written our equations the Bernoulli equation boundary condition we have written it up to order epsilon cube.

In particular some of the terms here this is a slightly lengthy exercise some of the terms are 0 even then we have 5 terms we are left with. We are trying to understand how these box terms

have a reason ok. So, once we understand these three the other two you can understand it in a similar manner.

So, I had written down one term which was like this and this would be in the Bernoulli equation in its primitive form before we did the expansion this would be $\frac{\partial \phi}{\partial t}$ at z is equal to η . And then I had done the transformation and I had written the perturbative expansion ok and then we had got up to here. Note that whatever I have written in the last line of the slide has come from the Taylor series expansion of just this term ok.

Now, let us understand where are these three green boxed expressions coming from ok. So, first the expression in which contains ω^2 , ok. So, that you can see very easily it is ω^2 into $\frac{\partial \phi_1}{\partial \tau}$, ok. So, ω^2 into $\frac{\partial \phi_1}{\partial \tau}$ the product of ϵ^2 into ϵ will make it an order ϵ^3 term ok. So, that is where this term is coming from.

What about this term η^2 into a third derivative of ϕ_1 ? So, η^2 will appear here there is a third derivative. So, the product of there will be an ϵ because there is a square here this will come out as ϵ^2 . So, the product of this ϵ^2 and that ϵ will make it an ϵ^3 . So, the product of this whole term with 1 here will just give me the term that I want ok. So, this term.

Similarly, you can find where we would have got this term from ok. We need an η^2 the only place here in this expression at the last line where η^2 appears is here. So, you can see that there is an η^2 here and then we need a second derivative with respect to ϕ_1 ok. So, that second derivative is this. So, it is a product of η^2 with this second derivative and once again with 1 ok.

So, it is there is an ϵ^2 here, there is an ϵ here, that makes in a order ϵ^3 cube and then if you take the product with 1 you will get this term. So, I hope it is clear how does one write these terms ok in order to determine ω^2 we will also have to write the modified Bernoulli equation.

The modified Bernoulli equation has an extremely lengthy right hand side and so, you if you follow this procedure you will be able to write all the terms that I am going to write in the modified Bernoulli equation next, ok.

But, before that let us make sure that we have understood how did we generate these terms on the right hand side of the Bernoulli equation boundary condition ok. So, this is how we got it. Now, you can also see there are other terms in if you look at the last line of this slide there are other terms which have appeared in this expression they are either epsilon to a smaller power. So, either they are order 1 or order epsilon square or maybe they are order epsilon 4.

If they are order epsilon square they would have already appeared in the previous equations, if they are order epsilon 4 they will appear in the next set of equations we are not going to write down equations up to order epsilon 4 it will not be necessary. So, I am only going to focus on from this last line we are only trying to understand which are the terms which appear at this order this is order epsilon cube.

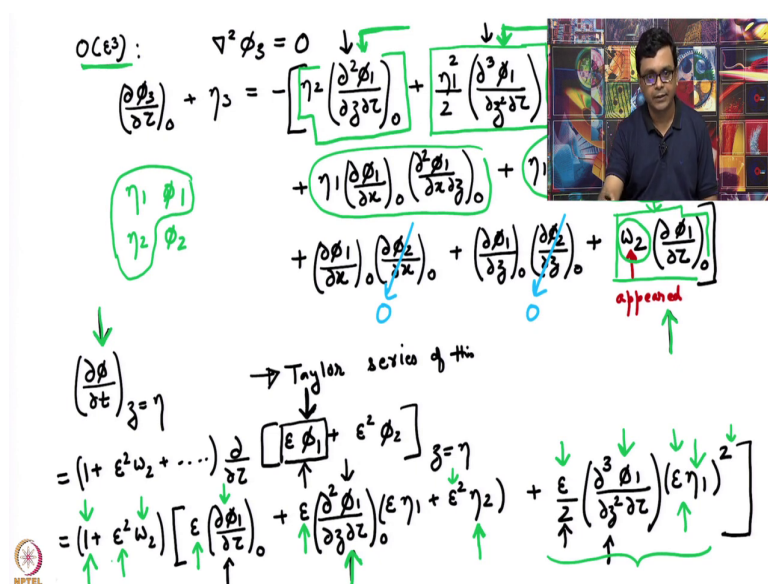
And so, we have explained now how these three box terms come from, in particular this is a term of interest because this term contains the reason why we had come to this order this contains ω^2 . And so, this is where these three boxes these three terms which are contained in the green boxes have come from ok they are all coming from this term in the Bernoulli equation.

Similarly, you can understand how does this term come from, how does this term come from go back to your primitive equations and try to understand which term gives birth to these terms. You will have to do a similar exercise keep writing Taylor series approximations, Taylor series expansions until you get all terms which are supposed to appear at a given order.

Note that even if you miss one term this is a very delicate step and you have to be very careful with your algebra, even if you miss one term your answers will be wrong unless that term is 0 or something like that your answers in general will be wrong. So, it is very important that one includes every term which is supposed to appear at a given order. So, now, let us proceed.

So, now we have determined the origin of these terms. Now, we need to do whatever we did at the previous order, we need to work out the functional form of all the terms which are not 0. Recall that the terms where there is a ϕ^2 appearing are all 0. So, I have put blue arrows indicating they are all 0 we do not have to worry about those terms, but we will have to work out the functional form of the remaining five terms, ok.

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The slide shows a handwritten derivation for the $O(\epsilon^3)$ term. At the top, it states $\nabla^2 \phi_3 = 0$. The main equation is:

$$\left(\frac{\partial \phi_3}{\partial t}\right)_0 + \gamma_3 = - \left[\gamma_2 \left(\frac{\partial^2 \phi_1}{\partial \partial^2 \partial t}\right)_0 + \frac{\gamma_1^2}{2} \left(\frac{\partial^3 \phi_1}{\partial \partial^2 \partial t}\right)_0 + \gamma_1 \left(\frac{\partial \phi_1}{\partial x}\right)_0 \left(\frac{\partial^2 \phi_1}{\partial x \partial \partial}\right)_0 + \gamma_1 \left(\frac{\partial \phi_1}{\partial x}\right)_0 \left(\frac{\partial^2 \phi_2}{\partial \partial^2}\right)_0 + \left(\frac{\partial \phi_1}{\partial x}\right)_0 \left(\frac{\partial \phi_2}{\partial x}\right)_0 + \left(\frac{\partial \phi_1}{\partial \partial}\right)_0 \left(\frac{\partial \phi_2}{\partial \partial}\right)_0 + \omega_2 \left(\frac{\partial \phi_1}{\partial t}\right)_0 \right]$$

Below this, a Taylor series expansion is shown for the time derivative term:

$$\left(\frac{\partial \phi}{\partial t}\right)_{\partial=\gamma} = (1 + \epsilon^2 \omega_2 + \dots) \left[\epsilon \left(\frac{\partial \phi_1}{\partial t}\right)_0 + \epsilon \left(\frac{\partial^2 \phi_1}{\partial \partial^2 \partial t}\right)_0 (\epsilon \gamma_1 + \epsilon^2 \gamma_2) + \frac{\epsilon}{2} \left(\frac{\partial^3 \phi_1}{\partial \partial^2 \partial t}\right)_0 (\epsilon \gamma_1)^2 \right]$$

The slide includes several annotations: green arrows pointing to specific terms, blue arrows pointing to terms that are zero, and a red arrow pointing to the ω_2 term labeled "appeared". A small video inset in the top right shows a lecturer.

If you do that you can do that because you know we know now η_1 , we know ϕ_1 , we know η_2 and we know ϕ_2 ; ϕ_2 has been used to set some terms to 0. So, you all you need to do is use the information of η_1 , η_2 and ϕ_1 to evaluate the derivatives that appear in these five terms and get the resultant expression on the right hand side. If you do that I have already done that I think you can do it yourself.

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$$\begin{aligned}
 \textcircled{B3} \quad \left(\frac{\partial \phi_3}{\partial \tau} \right)_0 + \eta_3 &= \frac{3}{8} \cos[3(x-\tau)] - \frac{3}{8} \cos(x-\tau) + \omega_2 \cos(x-\tau) \\
 \textcircled{C3} : \left(\frac{\partial^2 \phi_3}{\partial \tau^2} \right) + \left(\frac{\partial \phi_3}{\partial \eta} \right) &= - \left[\eta_2 \left(\frac{\partial^3 \phi_1}{\partial \eta \partial \tau^2} \right) + \frac{\eta_1^2}{2} \left(\frac{\partial^4 \phi_1}{\partial \eta^2 \partial \tau^2} \right) + \eta_1 \left(\frac{\partial^3 \phi_2}{\partial \eta \partial \tau^2} \right) \right. \\
 &\quad + \eta_2 \left(\frac{\partial^2 \phi_1}{\partial \eta^2} \right) + \frac{\eta_1^2}{2} \left(\frac{\partial^3 \phi_1}{\partial \eta^3} \right) + \eta_1 \left(\frac{\partial^2 \phi_2}{\partial \eta^2} \right) + 2\eta_1 \left(\frac{\partial \phi_1}{\partial \tau} \right) \left(\frac{\partial^3 \phi_1}{\partial \tau \partial \eta^2} \right) \\
 &\quad + 2\eta_1 \left(\frac{\partial^2 \phi_1}{\partial \tau \partial \eta} \right) \left(\frac{\partial^2 \phi_1}{\partial \tau \partial \eta} \right) + 2\eta_1 \left(\frac{\partial \phi_1}{\partial \eta} \right) \left(\frac{\partial^3 \phi_1}{\partial \tau \partial \eta^2} \right) \\
 &\quad + 2\eta_1 \left(\frac{\partial^2 \phi_1}{\partial \eta^2} \right) \left(\frac{\partial^2 \phi_1}{\partial \tau \partial \eta} \right) + 2 \left(\frac{\partial \phi_2}{\partial \tau} \right) \left(\frac{\partial^2 \phi_1}{\partial \tau \partial \eta} \right) \\
 &\quad + 2 \left(\frac{\partial \phi_2}{\partial \eta} \right) \left(\frac{\partial^2 \phi_1}{\partial \tau \partial \eta} \right) + 2 \left(\frac{\partial \phi_1}{\partial \eta} \right) \left(\frac{\partial^2 \phi_2}{\partial \tau \partial \eta} \right) + \frac{1}{2} \left(\frac{\partial \phi_1}{\partial \tau} \right) \frac{\partial}{\partial \tau} \left[\left(\frac{\partial \phi_1}{\partial \tau} \right)^2 + \left(\frac{\partial \phi_1}{\partial \eta} \right)^2 \right] \\
 &\quad \left. + \frac{1}{2} \left(\frac{\partial \phi_1}{\partial \eta} \right) \frac{\partial}{\partial \eta} \left[\left(\frac{\partial \phi_1}{\partial \tau} \right)^2 + \left(\frac{\partial \phi_1}{\partial \eta} \right)^2 \right] + 2\omega_2 \left(\frac{\partial^2 \phi_1}{\partial \tau^2} \right) \right]_0
 \end{aligned}$$

And so, I will give you the final answer; the final answer is $\frac{\partial \phi_3}{\partial \tau}$ at 0 plus η_3 after some amount of algebra is just $\frac{3}{8} \cos 3x \sin 3\tau - \frac{3}{8} \cos x \sin \tau + \omega_2 \cos x \sin \tau$. Note the appearance of the third harmonic now this is just an outcome of the fact that we are having cos cube, sin cube quantities this is at order epsilon cube and so, you will have get products which are cube of the primary.

The primary is $\cos(x - \tau)$ and so we are getting $\cos^3 \sin^3$ which can in turn can be expressed in terms of $\cos 3x \sin 3\tau$ and so on, ok. So, this is the third harmonic again a non-linear effect, ok. So, this is my equation B 3 with the functional form of the right hand side worked out.

In order to determine ω^2 we need to actually write down equation C 3 this is the lengthiest part of this exercise because in the equation for C 3 the number of terms on the right hand side is just too many.

So, I am just going to write down the equation C 3 and using the same argument that we have provided for equation B 3 you can see whether you can work out all the terms this is a very lengthy process. So, it will take me some time to write down all the terms which appear on the right hand side of C 3; C 3 is the modified Bernoulli boundary condition. So, C 3 is basically $\frac{\partial^2 \phi}{\partial \tau^2}$.

Now, I am going to write down all the terms. So, this term will be 0 because there is a ϕ^2 which appears here, this term will also be 0 because there is a ϕ^2 . There are total of 17 terms which appear on the right hand side. So, it is going to take some time to write all of them although it may seem that these terms are going to become very lengthy eventually there will be a lot of cancellation and the right hand side will become very simple.

Again this term is going to be 0 because it depends on ϕ^2 and all derivatives are applied at z is equal to 0 that term is also going to be 0 the previous one is also going to be 0. And then I have plus once again ω^2 makes its appearance in the last term and this whole thing gets applied at z is equal to 0.

So, every term gets evaluated all the derivatives of ϕ get evaluated at z is equal to 0. Clearly this is a very lengthy exercise, but it has to be done if we want to calculate what is the numerical value of ω^2 . In deep water it is the algebra is lengthy if we do this exercise in shallow water or even in finite depth the algebra is even lengthier. So, it is better to stick to the deep water approximation to get an idea of how does one determine ω^2 .

So, these are the terms and as I said earlier some of these will go to 0. So, this term will go to 0 there is a ϕ^2 there, this term will go to 0, then this term will go to 0 and this is also 0. So, total of 6 terms will go to 0. We are still left with 11 terms and I leave it to you to work out

the form of these terms you have to label them as 1 2 3 4 5 6 7 8 9 10 11 12 like that and then work out one by one, ok.

It is just a lengthy exercise it is not difficult because each of these terms we you know we know what is ϕ_1 , we know what is η_1 , we know ϕ_2 , η_2 . So, you just have to take the derivatives and it becomes easy to take the derivatives because the exponential factor does not really contribute you know sometimes it gets squared up it may become e to the power 2 z , but otherwise it just goes to 1.

So, you have to calculate this and the interesting part is when you have calculated all the non zero terms and then simplified them they will just there will be a lot of cancellations ok. So, there will be a lot of cancellations and you will find eventually that only one or two terms actually survive ok after doing all the cancellations you will see only one or two term survive. So, let me write down the final answer and you will see that the final answer is extremely compact.

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$$\left(\frac{\partial^2 \phi_3}{\partial \tau^2} + \frac{\partial \phi_3}{\partial \beta} \right)_0 = -\sin(x-\tau) + 2\omega_2 \sin(x-\tau)$$

$$= (2\omega_2 - 1) \sin(x-\tau) \quad \leftarrow \text{Resonant forcing term}$$

$\omega_2 = \frac{1}{2}$

$$\left(\frac{\partial^2 \phi_3}{\partial \tau^2} + \frac{\partial \phi_3}{\partial \beta} \right)_0 = \sin(x-\tau)$$

Soln to that eqn

P.I.: $\phi_3 = e^{\frac{1}{2}\tau} \cos(x-\tau)$

$$\left(\frac{\partial \phi_3}{\partial \tau} \right)_0 = \frac{1}{2} \cos(x-\tau) + \frac{1}{2} \tau \sin(x-\tau)$$

$$\left(\frac{\partial^2 \phi_3}{\partial \tau^2} \right)_0 = \sin(x-\tau) - \frac{1}{2} \tau \cos(x-\tau)$$

So, I am going to write down equation C 3 in its final form at z is equal to 0 or 0 let us write this straight line equal to minus sin x minus τ plus twice ω_2 sin x minus τ . So, all the eleven or twelve terms which survived this thing they internally canceled each other and we are left with only two terms both of which are proportional to sin of x minus τ . So, this is the simplification which happens ok.

So, now let me write this as twice ω_2 minus 1 sin of x minus τ . So, now, we need to determine ω_2 we do not want to solve the problem at order ϵ^3 we only are after the value of ω_2 . Now, in order to determine the value of ω_2 notice that this is a resonant forcing term. Why is this a resonant forcing term? So, I will show you that this is a resonant forcing term.

So, let us look at an equation whose form is this is equal to $\sin x \cos \tau$. I have just taken the left hand side of this and I have just replaced that by just unity you know I am just saying 1 into $\sin x \cos \tau$ this equation is of course, related to the equation written on top ok. I claim that the particular integral of this equation this is a differential equation it is a partial differential equation, but the left hand side is evaluated at z is equal to 0.

So, the particular integral of this equation is of the form $\frac{1}{2} \tau \cos x \sin \tau$ you can verify this we will verify this shortly, but you can see where. So, this is like our what we did, what we had in our ordinary differential equation case we had $t \sin t$ $t \cos t$ kind of terms which would come. Why is this coming? That is because this quantity $\sin x \cos \tau$ happens to solve this equation ok. So, happens to be a solution to the homogeneous equation, ok.

So, whenever you have something which is a solution to the homogeneous part then the particular integral has to be obtained by multiplying that by the time variable ok. So, we have this τ and for this particular right hand side there is a factor of half in the particular integral, you can check this very quickly that this actually solves the equation ok. So, this is a solution this is a solution to that equation solution to that equation.

Now, so you can see that $\frac{\partial \phi_3}{\partial \tau}$ would be basically $\frac{1}{2} \cos x \sin \tau$ we are basically saying ϕ_3 is ϕ_3 at 0 ok. So, not mentioned it ϕ_3 will have a yeah. So, this is this plus $\frac{1}{2} \tau \sin x \cos \tau$. And so, $\frac{\partial^2 \phi_3}{\partial \tau^2}$ is just $\sin x \cos \tau$ minus $\frac{1}{2} \cos x \sin \tau$. And if you work out the value of this derivative you know so, ϕ_3 will have a e to the power z . So, we should let me write e to the power z here let me get rid of the 0 here. So, that it is clear.

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Please note that the equation used here viz. $\left(\frac{\partial^2 \phi_3}{\partial \tau^2} + \frac{\partial \phi_3}{\partial \tau}\right)_{z=0} = \sin(x - \tau)$ is a model problem that demonstrates the resonant forcing that leads to secular terms. The ϕ_3 for this model problem is **not** the same as ϕ_3 for the Stokes wave in deep water.

$$\left(\frac{\partial^2 \phi_3}{\partial \tau^2} + \frac{\partial \phi_3}{\partial \tau}\right)_{z=0} = \sin(x - \tau)$$

$$\text{P.I.: } \phi_3 = e^{\frac{1}{2}\tau} \cdot \frac{1}{2} \tau \cos(x - \tau)$$

$$\left(\frac{\partial \phi_3}{\partial \tau}\right)_0 = \frac{1}{2} \cos(x - \tau) + \frac{1}{2} \tau \sin(x - \tau)$$

$$\left(\frac{\partial^2 \phi_3}{\partial \tau^2}\right)_0 = \sin(x - \tau) - \frac{1}{2} \tau \cos(x - \tau)$$

don't forget this

And if you work out $\frac{\partial \phi_3}{\partial \tau}$ by $\frac{\partial}{\partial \tau}$ at z is equal to 0 you will just get half $\tau \cos$ of x minus τ ok and if you add it to this part will get cancelled out and you will just be left with \sin of x minus τ . So, it is easy to check that this is a solution to this equation. What does this imply? This implies that every time we have something like this appearing on the right hand side anything proportional to \sin of x minus τ we are going to get a secular term appearing in our expression for ϕ .

If we; obviously, do not want that because we are looking at a wave problem, we are looking at travelling wave solutions. And so, the velocity potential cannot just grow in time ok if we want to eliminate that then we have to eliminate the term which causes this thing to appear.

We have looked at these kind of problems very early on in the course we are again encountering the same problem and we anticipating this problem we had already taken into account ω^2 . I encourage you to go and think that if we had not expanded time using the

Lindstedt Poincare scheme. What would have happened we would not have got a ω_2 and consequently we would have a resonant forcing term which could not be eliminated.

Now, we are going to eliminate the resonant forcing term and you can immediately see what is how to eliminate it we just have to set its coefficient to 0. So, this is just telling me that if I set ω_2 is equal to half then that problematic term. So, this is the problematic term which causes which causes this to appear ok. So, this will go away and the equation will become completely homogeneous.

Once again I remind you that we are not trying to solve the equation up to third order, we are just trying to determine ω_2 up to third order and our problem has been solved ω_2 has been determined and its numerical value is half. Let us now put all the things that we have done together ok.

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$$\begin{aligned}\eta &= \varepsilon \eta_1 + \varepsilon^2 \eta_2 + O(\varepsilon^3) \\ \phi &= \varepsilon \phi_1 + O(\varepsilon^3) \\ \eta &= \varepsilon \cos(x - \tau) + \frac{1}{2} \varepsilon^2 \cos\{2(x - \tau)\} + O(\varepsilon^3) \\ \phi &= \varepsilon e^{\frac{1}{2} \sin(x - \tau)} + O(\varepsilon^3) \\ k \hat{\eta} &= a_0 k \left[\cos\left\{k \hat{x} - \sqrt{gk} \hat{t} \left(1 + \frac{1}{2} a_0^2 k^2\right)\right\} \right. \\ &\quad \left. + \frac{1}{2} a_0 k \cos\left\{2\left(\frac{\quad}{\quad}\right)\right\} \right] + \dots \\ \Rightarrow \hat{\eta} &= a_0 \cos(k \hat{x} - \omega \hat{t}) + \frac{1}{2} a_0^2 k \cos[2(k \hat{x} - \omega \hat{t})] + \dots \\ \hat{\phi} &= \frac{a_0 g^{1/2}}{k^{1/2}} e^{k \hat{\phi}} \sin(k \hat{x} - \omega \hat{t})\end{aligned}$$

$\omega_2 = \frac{1}{2}$

$\omega = \sqrt{gk} \left(1 + \frac{1}{2} a_0^2 k^2 + \dots\right)$

$\boxed{\omega = \sqrt{gk}}$ non-linear connection

So, we have seen that η is $\epsilon \eta_1$ plus $\epsilon^2 \eta_2$ plus there would be order ϵ^3 corrections, ϕ is $\epsilon \phi_1$, ϕ_2 was 0. So, this would have an order ϵ^3 we anticipate that it could be ϵ^3 whether there is a ϕ_3 or not has to be actually gone back and we have to really solve the problem at order ϵ^3 in order to determine whether ϕ_3 is also 0 or not we have not done that.

So, that is why I am just writing order ϵ^3 it could be 0 then this would be order ϵ^4 . So, it is this and so, if I write η as $\epsilon \cos x$ minus τ plus half $\epsilon^2 \cos^2 x$ minus τ . I am just writing the value of η_2 plus order ϵ^3 , ϕ is just $\epsilon \sin x$ minus τ plus order ϵ^3 .

If we dimensionalize these expressions the utility of the fact that ω^2 is equal to half will become apparent if you dimensionalize these expressions. Let us dimensionalize them. So, I will write. So, all variables with hats are dimensional.

So, recall how we had non dimensionalize their expressions ϵ was a naught into k and so, this became \cos and the argument of \cos is $k x$ hat minus square root $g k$. Now, we are going back from τ to t to t hat. So, this is t hat into 1 plus half this is the half that is coming from ω^2 ok and ϵ^2 which is a naught square k^2 this is the argument of the first the primary mode ok plus half a naught square k^2 . And so, if I pull out the a naught into k then this just become a naught into k .

So, this is \cos again the same thing, but now with the 2 times you know plus dot dot dot. So, the whatever appears here is basically what I have written here. So, this is what it is ok . And so, this we can rewrite it as I can cancel out a k on both sides and I can rewrite this as \cos of $k x$ hat and now let me write this as ω into t hat ok and I will define what is ω plus half a naught square into $k \cos^2$ times the same thing plus dot dot dot.

Similarly, you can do the same exercise for ϕ ; ϕ is basically at this order there is no connection from non-linearity it is basically still the linear ϕ ok . So, we will just have a

naught g to the power half by k to the power half e to the power kz hat into $\sin kx$ hat minus ωt hat. And what have we defined as ω ?

Ω is basically just \sqrt{gk} into $1 + \frac{1}{2} a^2 k^2$ plus of course, there will be further corrections if we go to higher order. This is the important thing that we are finding this is the non-linear correction this is very reminiscent of what we had found for a non-linear pendulum.

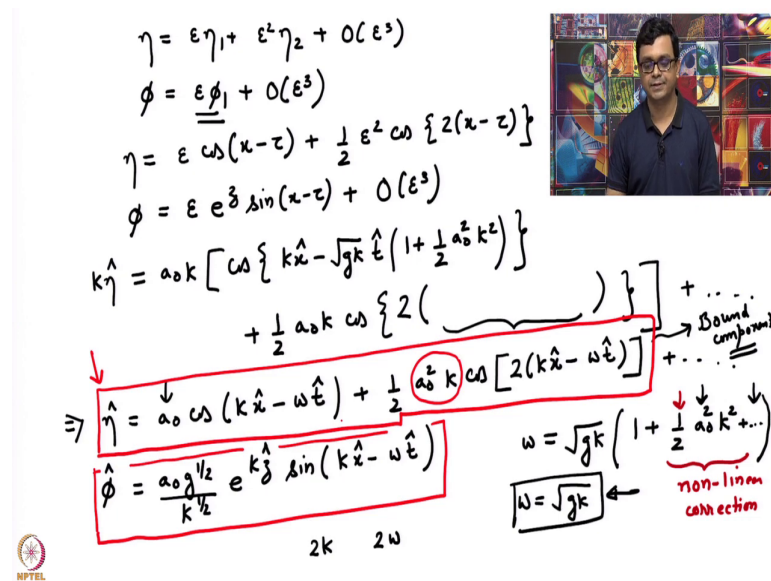
Now, we are doing a wave problem we have use the same Lindstedt Poincare technique this half is coming from there and we are finding you see this is at linear order if we had done it for sufficiently small a naught we would have just concluded ω is equal to \sqrt{gk} . This is just our deep water dispersion relation for surface gravity waves.

We are finding that what we had found earlier is an approximation the real dispersion relation is ω is equal to \sqrt{gk} into $1 + \frac{1}{2} a^2 k^2$ plus something and so on ok. We have we have to determine these terms we have to go to higher orders ok, but we have done the first correction ok.

So, this is telling us that the frequency for a given k for a given wavelength the frequency not only depends on the wavelength it also depends on the amplitude of the Fourier mode, the amplitude is a naught ok. So, the amplitude is a naught in addition there are corrections there are non-linear correction.

So, this is one correction that the frequency of the wave is different from that of the linear wave there is a correction to the dispersion relation there is an amplitude correction to the dispersion relation. In addition non-linearity introduces these higher harmonics we are do writing down up to order ϵ^2 . So, that is why the second harmonic has appeared.

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Handwritten mathematical derivations for the Stokes travelling wave solution:

$$\eta = \varepsilon \eta_1 + \varepsilon^2 \eta_2 + O(\varepsilon^3)$$

$$\phi = \varepsilon \phi_1 + O(\varepsilon^3)$$

$$\eta = \varepsilon \cos(x-z) + \frac{1}{2} \varepsilon^2 \cos\{2(x-z)\}$$

$$\phi = \varepsilon e^{\frac{1}{2} \sin(x-z)} + O(\varepsilon^3)$$

$$k \hat{\eta} = a_0 k \left[\cos\{k \hat{x} - \sqrt{g} k \hat{t} (1 + \frac{1}{2} a_0^2 k^2)\} \right. \\ \left. + \frac{1}{2} a_0 k \cos\{2(k \hat{x} - \omega \hat{t})\} + \dots \right]$$

Bound components

$$\Rightarrow \hat{\eta} = a_0 \cos(k \hat{x} - \omega \hat{t}) + \frac{1}{2} a_0^2 k \cos[2(k \hat{x} - \omega \hat{t})]$$

$$\hat{\phi} = \frac{a_0 g^{1/2}}{k^{1/2}} e^{\frac{1}{2} \sin(k \hat{x} - \omega \hat{t})}$$

$$\omega = \sqrt{g} k \left(1 + \frac{1}{2} a_0^2 k^2 + \dots \right)$$

Non-linear connection

NPTEL

Note that these are what are known as bound components. Why are they called bound components? Because notice that this mode has wave number $2k$ and a frequency 2ω you can check from this dispersion relation this is a non-linear dispersion relation ω is a non-linear function of k . So, if I make replace k by $2k$ the frequency will not become 2 times ω ok.

So, whatever is the frequency for k if I replace k by $2k$ the frequency will not double ok. And so, these components these bound components do not satisfy the dispersion relation. So, in this some sense they are bound to the primary mode ok.

So, this is basically the qualitative effect of non-linearity and this entire thing is basically what is known as the Stokes travelling wave solution. You can determine this up to more and

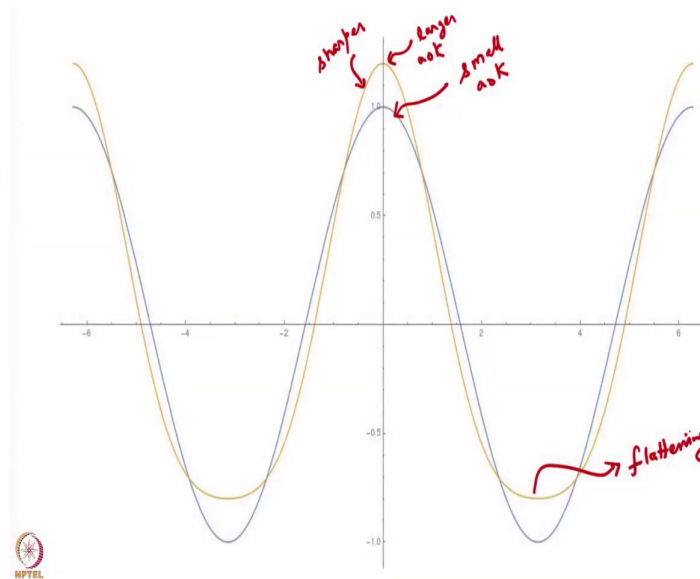
more orders it has already been done in the literature and Stokes travelling wave is known to a very high degree of accuracy.

Apart from the fact that we have discovered a non linear travelling wave solution which is very interesting this is also of practical use. It is used in many engineering contexts to model ocean waves in particular ocean wave breaking numerical simulations of ocean waves are initiated by modeling them as a Stokes wave as a large amplitude Stokes wave.

So, these are the qualitative features. Let us plot these waves up to the second harmonic and let us get a physical feel for what does the wave look like as we increase ϵ . So, you can see that the correction the if I put time t equal to 0, I can just plot η from this expression ok. So, I have done that and you can see I have done it for small value of ϵ and then larger value of ϵ ok.

So, for a small value of ϵ into k and a large value of ϵ into k ok. You will see that if ϵ for a given k if ϵ is sufficiently small then this second contribution will be very small ok and it will almost be a pure Fourier mode of the form $\cos kx$, but as ϵ gets larger and larger the second component makes a contribution.

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So, this is the picture this is small a naught into k . So, it is just a regular Fourier mode you know just a cosine I am plotting this a time t equal to 0. So, it is just $\cos kx$ here I have chosen k to be 1, but as I increase a naught this is larger a naught into k . Notice what is happening the crests are becoming sharper the troughs are flattening. So, there is a qualitative change in the shape of the wave there are lot of interesting things about these waves that have been under active investigation over many years now ok.

So, this is what we learn about the Stokes wave one can also do this exercise it is also of interest to derive the Stokes wave on a pool of finite depth the algebra gets even more lengthier. So, we are not going to do that here I will provide you a reference at the end of this video where you can read up about the expression for the Stokes wave the expression for η and for ϕ on for Stokes wave on a pool of finite depth a pool of finite depth h .

And there you will also find the correction to the dispersion relation the amplitude correction to the dispersion relation for a on a pool of finite depth, if you hold k constant and take h to

infinity you will recover the results that we have provided here. So, with this we come to the end of what we wanted to discuss on Stokes wave.

I hope you have seen now various aspects in this course we have done a very similar exercise using the Lindstedt Poincare technique earlier and we have found out the amplitude dependence on the frequency of the pendulum.

Now, in the last video we are finding a very similar thing for a surface gravity wave also. So, I hope you have understood the connection between these non-linear surface waves and the non-linear pendulum of course, the dynamics of these waves tend to be far more interesting there are interesting instabilities which happen for the Stokes wave so on and so forth and you can read some of those things on your own.

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For more on Stokes wave on finite depth, you may consult the following (see eqns. 2.382a,b c):

- Wave propagation on uneven bottoms, Dingemans, M.W., Doctoral Thesis. Downloadable freely at :

<http://resolver.tudelft.nl/uuid:67580088-62af-4c6f-b32e-b3940584e5d2>

